Table 1:	Columns	for I	P_meta	llicities.	_global.csv.
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Column	Unit	Description
name		Galaxy name
RA	$\operatorname{deg}$	Right ascension (J2000)
DEC	$\operatorname{deg}$	Declination (J2000)
Z		Heliocentric redshift
SF		Number of regions classified as starforming on BPT
AGN		Number of regions classified as AGN on BPT
Comp		Number of regions classified as composite on BPT
12+logOH_xx		Global 12+log(O/H) gas-phase metallicity using calibration xx
$12 + \log OH_xx_errdown$		16th percentile on distribution of 12+logOH_xx
12+logOH_xx_errup		84th percentile on distribution of 12+logOH_xx
$M_Z$ _xx	$M_{\odot}$	Mass of metals using metallicity calibration xx
$M_Z \_ xx\_ errdown$	$M_{\odot}$	16th percentile on distribution of $M_Z xx$
$M_Z \_ xx\_ errup$	$M_{\odot}$	84th percentile on distribution of $M_Z$ _xx
logNO_PG16		Global N/O ratio using calibration PG16
$\log NO_PG16_errdown$		16th percentile on distribution of logNO_PG16
$\log NO_PG16_errup$		84th percentile on distribution of logNO_PG16 $$

## Global metallicities for DustPedia galaxies

Radial profiles were fitted to the regions from DP\_metallicities\_regions.csv and global metallicities determined for 515 DustPedia galaxies by taking the metallicity at a radius of  $r = 0.4 r_{25}$ . These metallicities were derived using 6 different metallicity calibrations, yet we have opted to use the PG16S calibration to present our results as this calibration is more reliable at low metallicities. A bayesian approach was used to determine uncertainties on the gradients and global metallicities. An additional term of uncertainty from overlapping regions was added before fitting the gradients, and the Bayesian method also allows for intrinsic scatter. See Section 3.4 in De Vis et al. (2019) for more details. We also list the mass of metals for each calibration using  $M_Z = f_Z \times M_g + M_d$  where  $f_Z$  is the fraction of metals by mass calculated using  $f_Z = 27.36 \times 10^{(12+\log(O/H) - 12)}$ . The (total) gas masses  $M_g$  have been calculated using Equation 8 in Section 4.1 of De Vis et al. (2019). Uncertainties have been propagated using Monte Carlo bootstrapping.